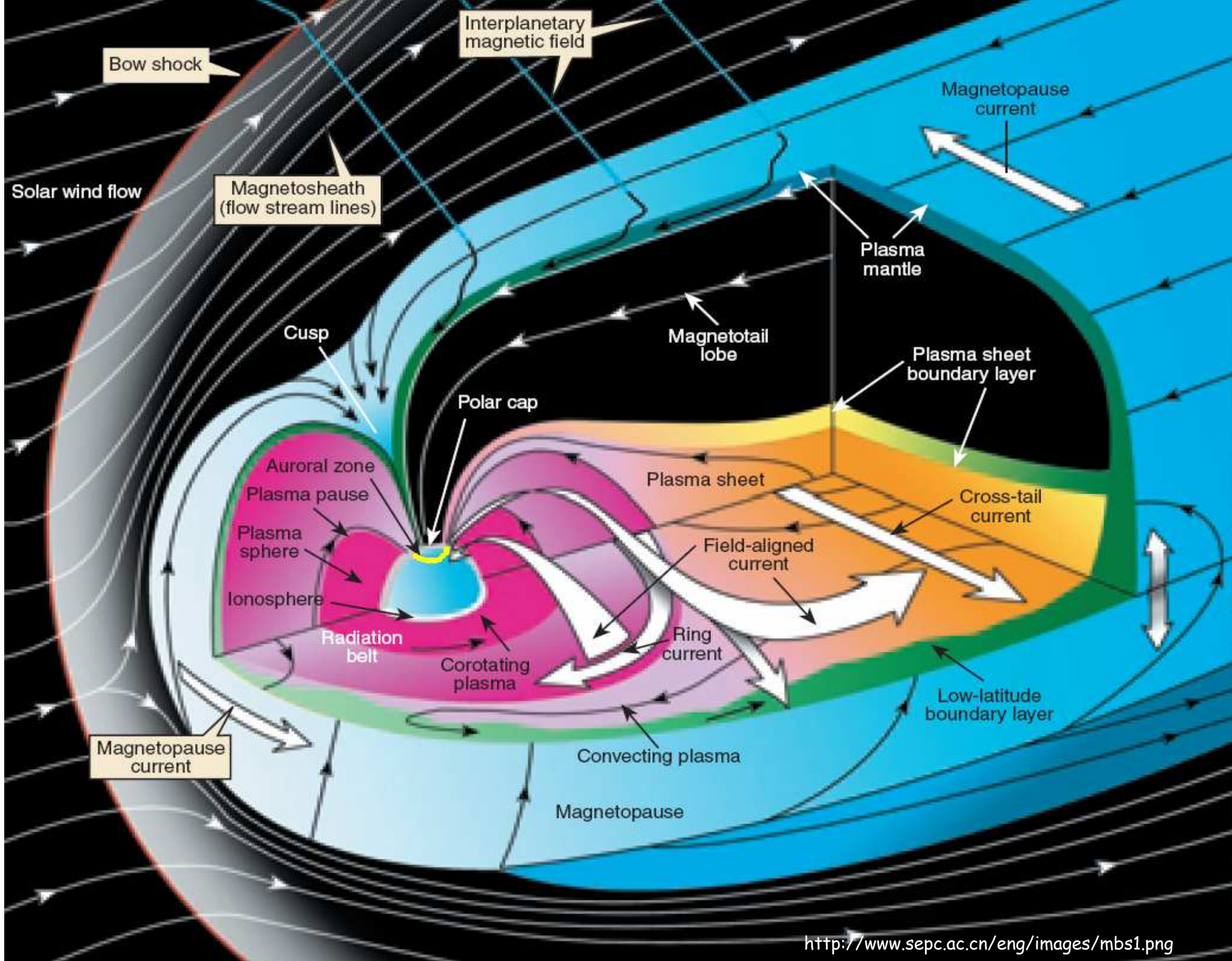
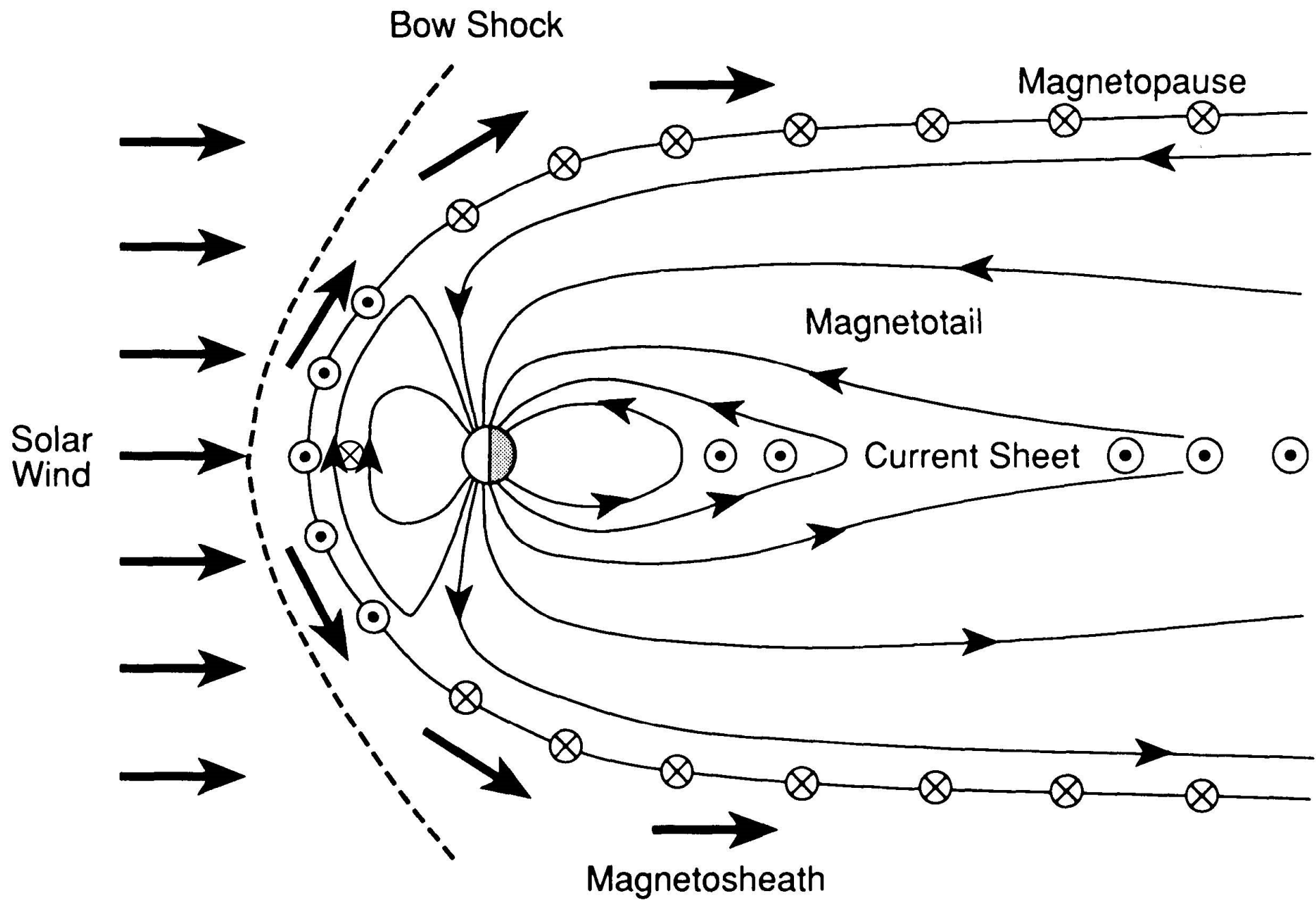






磁層子午面

Fig. 9.1 in M. G. Kivelson and C. T. Russell (1995)

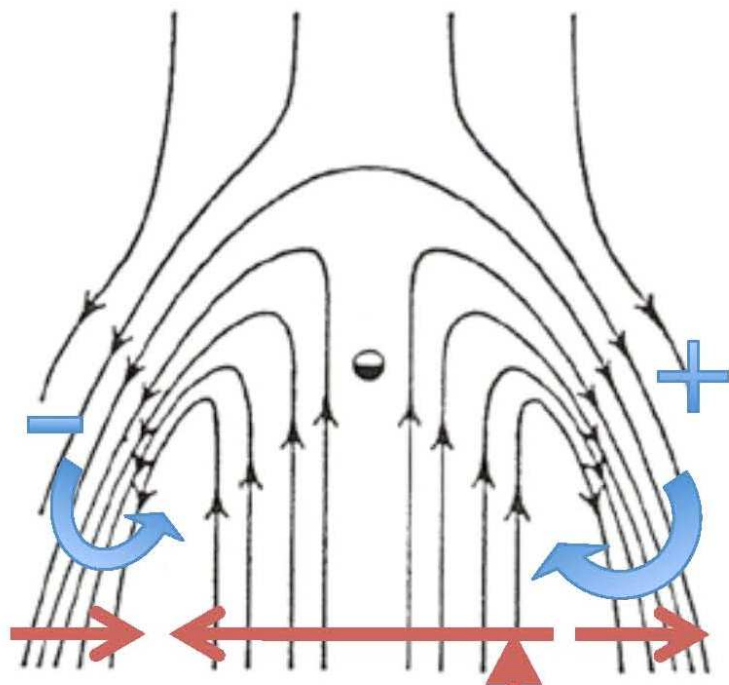
地球磁層赤道剖面圖（圖中標有箭頭的線，是流線）

太陽風吹過地球磁層

如果地球不自轉……

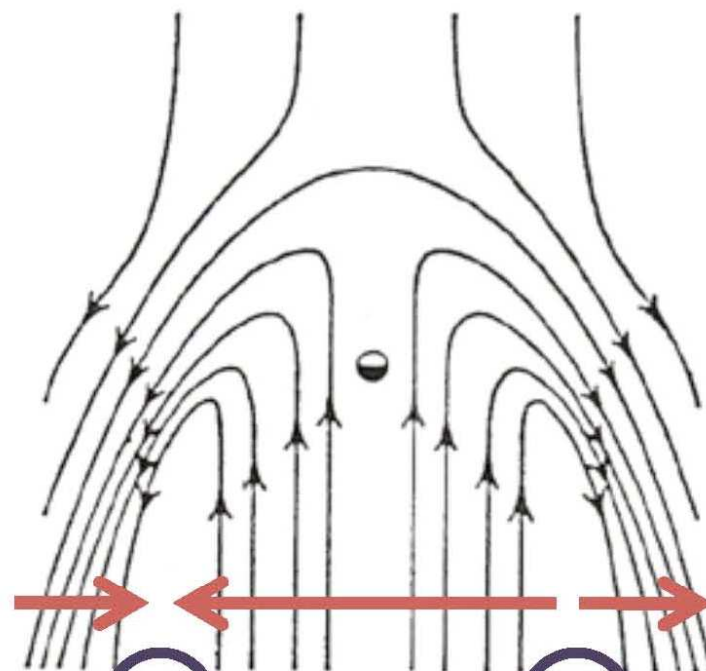


B



形成磁尾
電漿片與
南北磁尾腔

晨昏電場



Ω

$$-\mathbf{B} \cdot \Omega < 0$$

$$\nabla \cdot \mathbf{E}_{\perp} < 0$$



$$\Omega = \nabla \times \mathbf{V}$$

$$-\mathbf{B} \cdot \Omega > 0$$

$$\nabla \cdot \mathbf{E}_{\perp} > 0$$

$$\nabla \cdot \mathbf{E}_{\perp} = \nabla \cdot (-\mathbf{V} \times \mathbf{B})$$

$$= -\mathbf{B} \cdot (\nabla \times \mathbf{V}) + \mathbf{V} \cdot (\nabla \times \mathbf{B})$$

$$= -\mathbf{B} \cdot \Omega + \mathbf{V} \cdot \mathbf{J} / \mu_0$$

$$\approx -\mathbf{B} \cdot \Omega$$